4A15 Examiner's comments:

Q1 Oscillating sphere

The question was answered well by most candidates. It is similar to an example problem. The difference is that the boundary condition is on pressure instead of velocity. This confused some candidates.

Q2 Helmholtz resonator

A popular question. Most candidates did not take the effective length into account for part (b), and got the resonance frequency wrong by an order of magnitude. Almost all students got part (c) wrong. It turns out that the dimensions of the coke can are not really much smaller than the acoustic wavelength (an assumption that most candidates listed as an answer to part (a)), so the Helmholtz formula is not accurate here.

Q3 Multipole expansion

The least popular question, with a wide range of responses. Some were unfamiliar with multipole expansion. Some failed to realise that the strength of a dipole must be a vector. Several failed to use the correct Greens function for the Helmholtz equation, instead trying to use the Greens function for the wave equation. It is possible to derive the Greens function for the Helmholtz equation for the Helmholtz equation for the Helmholtz equation for the Helmholtz equation for the Helmholtz equation. It is possible to derive the Greens function for the Helmholtz equation for the Helmholtz equation for the Helmholtz equation for the Helmholtz equation for the Helmholtz equation.

Q4 Webster horn equation

A reasonably popular question. Some failed with bookwork to derive the Webster horn equation. Quite a few failed to spot that $0.5(1+\cos(2^*a^*x)) = \cos(a^*x)^2$, but succeeded otherwise with part (b). This affected their ability to come up with a simple answer for part (c), but the principle was mostly understood. Quite a few failed to notice that if alpha is taken to be imaginary, then the shape of the horn becomes $\cosh()$. (d) was mostly done correctly. Most candidates discussed the tuning of the harmonic series for a bugle in party (e)

Q1. Assuming that the pressure field
(A) bakes the form

$$P(x,t) = \frac{\partial}{\partial x} (f(r-cot))$$

We have assumed this form
because $f(r-cot)$ satisfies an
outgoing wave solution to the wave
equation. Therefore $\frac{\partial}{\partial x} (f(r-cot))$
must also be a solution to the
wave equation.
Let $f(r-cot) = de co$
 $= e^{i}(wt-kr)$
 $p'(x,t) = \frac{\partial}{\partial x} f(e^{i}(wt-xr))$
Neve it is assumed that the

physical pressure field will be the real part of this expression. $(k = w/c_0)$ $b((r, 0, t) = \begin{cases} -1 & -ik \\ r^2 & t \end{cases}$ $k \in r$ $\chi \in r$



The BC at r=a can be written as b'(r=a,0,t)= Re f A cuso e^{iwt} -(2) Substituting r=a in Eq(1) ad equation it to be RMS in (2):

$$\alpha \left\{ \begin{array}{c} -\frac{1}{a^{2}} - \frac{i\kappa}{a} \right\} e^{i(\omega t - \kappa a)}$$

$$= A \cos e^{i\omega t}$$

$$\Rightarrow A \cos e^{i\omega t}$$

$$\Rightarrow \alpha = -A e^{i\kappa a} a^{2} - (5)$$

$$(1 + i\kappa a)$$

It is easier to calculate the
radiated power in the farfield,
In the far field,
$$\frac{P'}{u'} \sim Poco$$
.
 $\frac{P'}{u'} = \frac{P'}{Poco} - \frac{14}{2}$
Also, as $r \rightarrow \infty$,





$$= \frac{2}{3}$$

$$Eq.(6) becomes$$
Power = $\frac{\chi^2}{3} \frac{\beta t^2 2TT}{\beta co}$
Using Eq.(3),
Power = $2TT \frac{k^2}{3} \frac{A^2 a^4}{\beta co} (1 + k^2 a^2)$
Power = $2TT \frac{a^2 A^2}{3} \frac{k^2 a^4}{\beta co} (1 + k^2 a^2)$

$$Power = 2TT \frac{a^2 A^2}{3} \frac{k^2 a^4}{\beta co} (1 + k^2 a^2)$$

$$-(9)$$

ka << 1 (compact source) & X <La >> Ka>>1 Form (9). Power (rea. >>1) ~ Ztt arAr Ztt Poco & Power (Kacc) $\sim \frac{2}{3} + \frac{a^2 A^2 (k^2 a)^2}{\rho_{olo}}$ Confiring these two exposing we see that to get the same power A needs to be higher by a factor of (ka) when

When Kall to get the Same prover when Ka >>1. Eg J ka = 0:(, A needs to be 10 tons hager to rasole the same pour as When ka = 10. QZ (A) . Narrow spening connected to a large cevily · Small dimensions of the neck & Carily relative to the acoustic wavelength · Viscous danping is regligible · Perturbations are linear and isentropic $A = 4 \times 10^{-4} m^2$ (6) $\pi a^2 = A \implies a = \sqrt{A} = 1.1 \times 10^2 M$ The effective length of the neck L'N. 1+ 1.20 N1.20 (1<<a) $V = 0.33 L = 0.33 \times 10^3 M^3$ $f = \frac{6}{2\pi} \int \frac{A}{iv} = \frac{347}{24} \int \frac{4x10^{-4}}{2x14x10^{-2} \times 0.33x10^{-3}}$ f = 529.2 Hz

(c) For
$$f = 529.2 \text{ M}_2$$
, the wavelength
 $\lambda = \frac{347}{529.2} = 0.66 \text{ M}_2$

The con his a height of DilM. The wavelength is not much larger that the size of the cavity. -Fix a given & wavelength, a lorger Cavity has a lower shiftness company to a smaller cavity. The lower shiftness work result in a lower resonance forces



$$f = \frac{c}{\chi} = \frac{347}{54} = \frac{867.5 \text{ Hz}}{54}$$

$$\frac{M_{n}H_{ip}de}{p} = \frac{2}{n} \frac{S_{n}}{S_{n}} \frac{F(n - n)}{F(n - n)} = \frac{1}{p} e^{-ik(n)}$$

$$G = \left(\frac{1}{2} \frac{S_{n}}{S_{n}}\right) = \frac{1}{p} \left(\frac{n}{2} - \frac{1}{2} \frac{1}{n} + \frac{1}{2} \frac{1}{n}\right) = \frac{1}{p} \left(\frac{1}{2} \frac{S_{n}}{S_{n}}\right) = \frac{1}{p} \left(\frac{1}{2} \frac{S_{n}}{S_{n}} \frac{1}{2} \frac{1}{$$

c)
$$\mathcal{L}S_n n_n : \begin{pmatrix} 3 \\ 0 \end{pmatrix} + \begin{pmatrix} 3 \\ 0 \end{pmatrix} + \begin{pmatrix} 0 \\ 1 \end{pmatrix} \begin{pmatrix} 0 \\ 1 \end{pmatrix} \begin{pmatrix} 0 \\ 1 \end{pmatrix}$$

 $\frac{1}{5} \begin{pmatrix} 6 \\ 10 \end{pmatrix}$
He significance of Ligher order terms in the
multipole expansion increases as frequency
hoer, so using just the dipole term will broad
less accurate of the frequency increases.

$$d) p = \int \frac{e^{-ik|n-y|}}{4\pi \int y} 4\pi \int y dy$$
$$= \int \frac{e^{-ik|n-y|}}{|n-y|} \int y dy$$
$$= \int f(n-y) \int y dy$$

e) $f(n-y) = f(n) - y \cdot \nabla f(n) + ...$ $\beta = \left(\int S(y) dy \right) f(n) - \left(\int S(y) y dy \right) \cdot \nabla f(n)$ A t. J voorde dopole monopole

reboter Harn Questio G) (**ð** CV 0 AN 3 3 Sn n (\mathcal{P})

$$\int_{CV} \frac{1}{2} \int_{CV} \frac{1}{2} \frac{1}{2$$

on
$$\mathbb{D}$$
 \mathbb{E} \mathbb{E} n . $\nabla p = 0$ bucke of
the would
on \mathbb{D} n . $\nabla p = -\frac{2t}{2n}$
on \mathbb{D} n . $\nabla p = -\frac{2t}{2n}$
on \mathbb{D} n . $\nabla p = \frac{2t}{2n}$
asking p , $\frac{2t}{2n}$, one constant
over cross-section
 $\int \frac{1}{c} \frac{3t}{2t} N + \frac{3t}{2n} \Big|_{n} A_{n} - \frac{3t}{2n} \Big|_{n \in \mathbb{N}} A_{(n+S_n)} = 0$
let $S_n = 0$
 $\frac{A S_n}{c} \frac{3t}{2t^2} - S_n \frac{3}{2n} \Big(A \frac{3t}{2n}\Big) = 0$

$$\frac{1}{2} \frac{3^2}{3^2} - \frac{1}{4} \frac{3}{2} \left(A \frac{3}{2^2} \right) = 0$$

b)
$$A = \frac{A_0}{2} \left(1 + \cos(2\alpha n) \right)$$

 $= A_0 \cos^2(\alpha n) = \pi \alpha^2$
 $a = \int_{\overline{M}}^{\overline{M}_0} \cos(\alpha n)$
 $\frac{da}{dn} = -\int_{\overline{M}}^{\overline{M}_0} \alpha \sin(\alpha n)$
 $\frac{da}{dn^2} = -\int_{\overline{M}}^{\overline{M}_0} \alpha^2 \cos(\alpha n)$
 $\frac{1}{\alpha} \frac{d^2\alpha}{dn^2} = -\int_{\overline{M}}^{\overline{M}_0} \alpha^2 \cdot \int_{\overline{M}_0}^{\overline{\pi}}$
 $= -\alpha^2$
 $A_{U} \left(-\alpha \right)$

$$\frac{d^2 \varphi}{dn^2} + \left(k^2 + \alpha^2\right) \varphi = O$$

$$\begin{aligned} \varphi &= I e^{-i\sqrt{k^{2}+k^{2}}n} + Re^{i\sqrt{k^{2}+k^{2}}n} \\ \hat{\rho} &= A^{-\frac{1}{2}} \psi \\ &= A^{-\frac{1}{2}} \frac{\varphi^{-1}(\varphi n)}{\cos^{-1}(\varphi n)} \psi \\ &= A^{-\frac{1}{2}} I \frac{e^{-i\sqrt{k^{2}+k^{2}}n}}{\cos(\varphi n)} + A^{-\frac{1}{2}} R \frac{e^{i\sqrt{k^{2}+k^{2}}n}}{\cos(\varphi n)} \end{aligned}$$

() for no propagation me need $l_{\mu}^{+} \alpha^{2} \leq 0$ take a to be imaginery aziB k-BCO B> k if a is weplex the the solins becomes a cosh and $a[n] = \int_{\overline{D}}^{\overline{R}_{0}} con(\frac{R_{0}}{2}) = \int_{\overline{D}}^{\overline{R}_{0}} cosh(\frac{3n}{2})$ $A_{b_1} = A_0 \cosh\left(\frac{B_n}{2}\right)$



d) i) $\frac{1}{5} \frac{d^2 G}{dn^2} = O$ =) Il plane comes propagate \vec{r}) $(dl_{n} = 0)$ > M place mens propagate.

e) brok Lorns. The bell is designed so that I déa increases mening louer frequencies replict deponside HI Lell. This helps the bugle be able to produce a full hornonic series despite having closed - open boundary condifiens.

Deriction of Modified NOT IN QUESTION.

$$k^{2} \hat{\rho} + \frac{1}{A} \frac{d}{dn} \left(\frac{d \hat{\rho}}{dn} \right) = 0$$

$$\psi(n) = \hat{\rho}(n) An'^{h}$$

$$\hat{\rho} = \psi A^{-h} \qquad A = \pi n^{h}$$

$$(n - h - 1)$$

$$= \Psi \pi a$$

$$\frac{l^2 \Psi}{a} \Psi \frac{l}{b^2} \frac{d}{dt} \left(\frac{d}{dt} \left(\frac{d}{dt} \left(\frac{\psi}{a} - \frac{d}{dt} \right) \right) \right) = 0$$

$$\frac{k^{2}}{a}\psi + \frac{1}{a^{2}}\frac{d}{dn}\left(a^{2}\left(\frac{1}{a}\frac{dY}{dn} - \frac{\psi}{a^{2}}\frac{da}{dn}\right)\right) = 0$$

$$\frac{k^{2}}{a}\psi + \frac{1}{a^{2}}\frac{d}{dn}\left(a\frac{d\Psi}{dn} - \frac{\psi}{dn}\frac{da}{dn}\right) = 0$$

$$\frac{k^{2}}{a}\psi + \frac{1}{a^{2}}\left(c\frac{d^{2}\psi}{dn^{2}} + \frac{da}{dn}\frac{d\Psi}{dn} - \frac{d\psi}{dn}\frac{da}{dn} - \frac{\psi}{dn^{2}}\frac{da}{dn^{2}}\right) = 0$$

$$\frac{k^{2}}{a}\psi + \frac{d^{2}\psi}{dn^{2}} - \frac{\psi}{a}\frac{d^{2}a}{dn} = 0$$

$$\frac{d^{2}\psi}{dn^{2}} + \frac{d^{2}\psi}{dn^{2}} - \frac{\psi}{a}\frac{d^{2}a}{dn} = 0$$

$$\frac{d^{2}\psi}{dn^{2}} + \left(k^{2} - \frac{1}{a}\frac{d^{2}a}{dn^{2}}\right) = 0$$