EGT3
ENGINEERING TRIPOS PART IIB

Monday 23 April $2018 \quad 9.30$ to 11.10

## Module 4A15

## AEROACOUSTICS

Answer not more than three questions.
All questions carry the same number of marks.
The approximate percentage of marks allocated to each part of a question is indicated in the right margin.

Write your candidate number not your name on the cover sheet.

## STATIONERY REQUIREMENTS

Single-sided script paper

## SPECIAL REQUIREMENTS TO BE SUPPLIED FOR THIS EXAM <br> Engineering Data Book <br> CUED approved calculator allowed <br> Attachment: 4A15 Aeroacoustics data sheet (6 pages).

10 minutes reading time is allowed for this paper at the start of the exam.

You may not start to read the questions printed on the subsequent pages of this question paper until instructed to do so.

## Version AA/5

1 An earthquake can result in surface waves of significant amplitude. These waves radiate sound into the atmosphere. As a result of the Sumatran earthquake in 2004, surface waves with a vertical velocity amplitude of $1 \mathrm{~cm} \mathrm{~s}^{-1}$ and a time period of 20 s were recorded. The speed of the surface wave was eight times the speed of sound in air. In the following, assume that the Earth is flat and the surface waves are plane and have a constant amplitude.
(a) Determine the direction of propagation of the plane sound wave.
(b) Determine the mean intensity of the plane waves radiated into the atmosphere. Assume, at sea level, the atmosphere has a mean density of $\rho_{0}=1.2 \mathrm{~kg} \mathrm{~m}^{-3}$ and speed of sound of $c_{0}=340 \mathrm{~m} \mathrm{~s}^{-1}$.
(c) Seismic activity can be detected in space. By applying conservation of energy of the sound wave and neglecting sound dissipation, refraction and reflection for these low-frequency waves, find the velocity amplitude induced by the sound wave in the ionosphere. Assume that the gas density in the ionosphere is $10^{-8} \rho_{0}$. Because the speed of sound does not vary as drastically as the density with altitude, assume that $c_{0}$ is a constant for the purpose of getting an order of magnitude estimate. You may also assume that the spatial extent of the surface wave is much longer than the distance from the surface of the Earth to the ionosphere, so that the waves remain plane with no spherical spreading.

## Version AA/5

2 Free reeds, found in musical instruments such as the harmonica, produce a tone when air is sucked through them because the flow makes the reed oscillate within the reed block at its resonant frequency. This modulates the opening of the reed resulting in a pulsating mass flow. In the following assume that the unsteady mass flow per unit volume at the reed is $\dot{\mu}(\mathbf{x}, t)$, where $\mathbf{x}$ is the position and $t$ is the time (a dot over a letter denotes a time derivative).
(a) Show that the equation governing sound propagation from the reed is given by

$$
\frac{\partial^{2} \rho^{\prime}}{\partial t^{2}}-c_{0}^{2} \nabla^{2} \rho^{\prime}=\ddot{\mu}(\mathbf{x}, t)
$$

where $\rho^{\prime}$ is the acoustic density perturbation and $c_{0}$ is the ambient speed of sound.
(b) If the unsteady mass flow through the reed is

$$
\dot{m}=\int_{V} \dot{\mu}(\mathbf{y}, t) d \mathbf{y}
$$

where $V$ is the source volume, then assuming that the reed is a spatially compact source radiating sound into free space, show that the acoustic pressure field radiated at a large distance $r$ from the reed is given by

$$
p^{\prime}(r, t)=\frac{\ddot{m}\left(t-r / c_{0}\right)}{4 \pi r}
$$

(c) Assuming $\dot{m}(t)=\rho_{0} Q[1+\cos (2 \pi f t)]$, where $\rho_{0}=1.2 \mathrm{kgm}^{-3}$ is the ambient density of air and $Q=12 \mathrm{~L} / \mathrm{min}$ is the flowrate, $f=261.6 \mathrm{~Hz}$ is the resonant frequency of the reed, calculate the SPL of the sound radiated at a distance of 10 m from the reed.
(d) If, instead of sucking air through the reed, the reed is mechanically plucked it makes very little sound. Explain why this is the case.

## Version AA/5

3 Acoustic pressure fluctuations, $p^{\prime}$, in a fluid which has a uniform mean flow in the $x_{3}$-direction at a Mach number $M$ satisfy the convective wave equation

$$
\nabla^{2} p^{\prime}-\left(\frac{1}{c_{0}^{2}} \frac{\partial^{2} p^{\prime}}{\partial t^{2}}+\frac{2 M}{c_{0}} \frac{\partial^{2} p^{\prime}}{\partial t \partial x_{3}}+M^{2} \frac{\partial^{2} p^{\prime}}{\partial x_{3}^{2}}\right)=0
$$

where $c_{0}$ is the speed of sound.
A rigid walled duct of rectangular cross-section has a mean flow through it with Mach number $M$ in the $x_{3}$-direction. The duct has width $a$ and height $b$. The coordinate system is shown in Fig 1.
(a) Explain why the pressure perturbation in acoustic modes of frequency $\omega$ within the duct can be written in the form

$$
p^{\prime}(\mathbf{x}, t)=e^{i \omega t} g\left(x_{3}\right) \cos \left(\frac{m \pi x_{1}}{a}\right) \cos \left(\frac{n \pi x_{2}}{b}\right)
$$

for some function $g\left(x_{3}\right)$ and integer constants $m$ and $n$. Detailed calculations are not required.
(b) By trying a solution of the form $g\left(x_{3}\right)=A \exp \left(i k x_{3}\right)$, where $A$ and $k$ are constants, find the function $g\left(x_{3}\right)$ in terms of $k_{0}\left(=\omega / c_{0}\right)$ and $M$.
(c) Interpret your solution for $g\left(x_{3}\right)$ when $m=n=0$.
(d) For a mode with non-zero $m$ and/or $n$, determine the frequency range for which that mode is cut-off.


Fig. 1

## Version AA/5

4 Consider a wall-bounded two-layer fluid system, in which the region $x<0$ contains a fluid of density $\rho_{0}$ and sound speed $c_{0}$, the region $0<x<h$ contains a fluid of density $\rho_{1}$ and sound speed $c_{1}$, and an infinite wall with complex impedance $Z$ is located at $x=h$. In $x<0$ an incident plane sound wave of amplitude $I$ and frequency $\omega$ propagates parallel to the $x$ axis.
(a) Show that the reflected wave in $x<0$ has complex reflection coefficient

$$
R=I \frac{\left(1-Z^{*}\right)\left(1+\rho^{*}\right)+\left(Z^{*}+1\right)\left(1-\rho^{*}\right) \exp \left(2 \mathrm{i} k_{1} h\right)}{\left(1-Z^{*}\right)\left(\rho^{*}-1\right)-\left(Z^{*}+1\right)\left(1+\rho^{*}\right) \exp \left(2 \mathrm{i} k_{1} h\right)}
$$

where

$$
\rho^{*}=\frac{c_{1} \rho_{1}}{c_{0} \rho_{0}}, Z^{*}=\frac{Z}{c_{1} \rho_{1}}
$$

and $k_{1}=\omega / c_{1}$.
(b) Calculate the modulus and phase of $R / I$ when $Z^{*}=1$. Comment on the physical implication of your results.

## END OF PAPER

Version AA/5

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## Module 4A15 Aeroacoutics Data Sheet

## USEFUL DATA AND DEFINITIONS

## Physical Properties

Speed of sound in an ideal gas $\sqrt{\gamma R T}$, where $T$ is temperature in Kelvins

## Units of sound measurement

$$
\begin{aligned}
\text { SPL (sound pressure level) } & =20 \log _{10}\left(\frac{p_{r m s}^{\prime}}{2 \times 10^{-5} \mathrm{Nm}^{-2}}\right) \mathrm{dB} \\
\text { IL (intensity level) } & =10 \log _{10}\left(\frac{\text { intensity }}{10^{-12} \text { watts } \mathrm{m}^{-2}}\right) \mathrm{dB} \\
\text { PWL (power level) } & =10 \log _{10}\left(\frac{\text { sound power output }}{10^{-12} \text { watts }}\right) \mathrm{dB}
\end{aligned}
$$

## Definitions

Surface impedance $Z_{s}$, relates the pressure perturbation applied to a surface, $p^{\prime}$, to its normal velocity $v^{\prime} ; p^{\prime}=Z_{s} v^{\prime}$

Characteristic impedance of a fluid $\rho_{0} c_{0}$
Specific impedance of a surface $Z_{s} /\left(\rho_{0} c_{0}\right)$
Wavenumber $k=\omega / c_{0}=2 \pi / \lambda$, where $\lambda$ is the wavelength
Helmholtz number (or compactness ratio) $=k D$, where $D$ is a typical dimension of the source.

Strouhal number $=\omega D /(2 \pi U)$ for sound of frequency $\omega($ in $\mathrm{rad} / \mathrm{s})$, produced in a flow with speed $U$, length scale $D$.

## Basic equations for linear acoustics

## Conservation of mass

$$
\frac{\partial \rho^{\prime}}{\partial t}+\rho_{0} \nabla \cdot \mathbf{v}^{\prime}=0
$$

## Conservation of momentum

$$
\rho_{0} \frac{\partial \mathbf{v}^{\prime}}{\partial t}+\nabla p^{\prime}=0
$$

## Isentropic

$$
c_{0}^{2}=\left.\frac{d p}{d \rho}\right|_{S}
$$

## Wave equation

$$
\frac{1}{c_{0}^{2}} \frac{\partial^{2} p^{\prime}}{\partial t^{2}}-\nabla^{2} p^{\prime}=0
$$

## Energy density

$$
e=\frac{1}{2} \rho_{0} v^{\prime 2}+\frac{1}{2 \rho_{0} c_{0}^{2}} p^{\prime 2}
$$

Intensity $\mathbf{I}=p^{\prime} \mathbf{v}^{\prime}$
Velocity potential $\phi^{\prime}$ satisfies the wave equation and $p^{\prime}=-\rho_{0} \frac{\partial \phi^{\prime}}{\partial t}, \mathbf{v}^{\prime}=\nabla \phi^{\prime}$.
Autocorrelation $F(\xi)$, the autocorrelation of $f(y)$ is given by

$$
\begin{gathered}
F(\xi)=\overline{f(y) f(y+\xi)} \\
F(0)=\overline{f^{2}}
\end{gathered}
$$

## Integral length scale, $l$

$$
l \overline{f^{2}}=\int_{-\infty}^{\infty} F(\xi) d \xi
$$

## Sound power

Sound power from a source is defined as

$$
P=\int_{S} \overline{\mathbf{I}} \cdot \mathbf{d S}=\int_{S_{\infty}} \frac{\overline{p^{\prime 2}}}{\rho_{0} c_{0}} \mathbf{d S}
$$

for a statistically stationary source. For an outward propagating spherically symmetrical sound field $P=\frac{\overline{p^{\prime 2}}}{\rho_{0} c_{0}} 4 \pi r^{2}$, where $p^{\prime}$ is the acosutic pressure at radius $r$.

For a sound field, which is a function of spherical polar coordinates $r, \boldsymbol{\theta}$ only, and is independent of the azimuthal angle,

$$
P=2 \pi r^{2} \int_{0}^{\pi} \frac{\overline{p^{\prime 2}}}{\rho_{0} c_{0}} \sin \theta d \theta
$$

## Simple wave fields

## 1D or plane wave

The general solution of the 1D wave equation is $p^{\prime}(x, t)=f\left(t-x / c_{0}\right)+g(t+$ $x / c_{0}$ ), where $f$ and $g$ are arbitrary functions. In a plane wave propagating to the right $p^{\prime}=\rho_{0} c_{0} u^{\prime}$; in a plane wave propagating to the left $p^{\prime}=-\rho_{0} c_{0} u^{\prime}, u^{\prime}$ being the particle velocity.

## Spherically symmetric sound fields

The general spherically symmetric solution of the 3D wave equation is

$$
\phi^{\prime}(r, t)=\frac{f\left(t-r / c_{0}\right)}{r}+\frac{g\left(t+r / c_{0}\right)}{r}
$$

where $r$ is the distance from the source; $f$ and $g$ are arbitrary functions.

## $\cos \theta$ dependence

The general solution of the 3 D wave equation with $\cos \theta$ dependence is

$$
p^{\prime}(\mathbf{x}, t)=\frac{\partial}{\partial x}\left[\frac{f\left(t-r / c_{0}\right)}{r}+\frac{g\left(t+r / c_{0}\right)}{r}\right]=\cos \theta \frac{\partial}{\partial r}\left[\frac{f\left(t-r / c_{0}\right)}{r}+\frac{g\left(t+r / c_{0}\right)}{r}\right]
$$

## Useful mathematical formulae

Spherical polar coordinates $(r, \theta, \psi)$

## Gradient

$$
\nabla p^{\prime}=\left(\frac{\partial p^{\prime}}{\partial r}, \frac{1}{r} \frac{\partial p^{\prime}}{\partial \theta}, \frac{1}{r \sin \theta} \frac{\partial p^{\prime}}{\partial \psi}\right)
$$

## Divergence

$$
\nabla \cdot \mathbf{v}^{\prime}=\frac{1}{r^{2}} \frac{\partial}{\partial r}\left(r^{2} v_{r}^{\prime}\right)+\frac{1}{r \sin \theta} \frac{\partial}{\partial \theta}\left(\sin \theta v_{\theta}^{\prime}\right)+\frac{1}{r \sin \theta} \frac{\partial v_{\phi}^{\prime}}{\partial \psi}
$$

## Laplacian

$$
\nabla^{2} p^{\prime}=\frac{1}{r^{2}} \frac{\partial}{\partial r}\left(r^{2} \frac{\partial p^{\prime}}{\partial r}\right)+\frac{1}{r^{2} \sin \theta} \frac{\partial}{\partial \theta}\left(\sin \theta \frac{\partial p^{\prime}}{\partial \theta}\right)+\frac{1}{r^{2} \sin ^{2} \theta} \frac{\partial^{2} p^{\prime}}{\partial \psi^{2}}
$$

## Delta functions

## Kronecker Delta

$$
\delta_{i j}= \begin{cases}1 & i=j \\ 0 & i \neq j\end{cases}
$$

1D $\delta$-function $\delta(x)=0$ for $x \neq 0$ and $\int_{-\infty}^{\infty} \delta(a x-b) f(x) d x=f(b / a) /|a|$
3D $\delta$-function $\delta(\mathbf{x})=\boldsymbol{\delta}\left(x_{1}\right) \boldsymbol{\delta}\left(x_{2}\right) \boldsymbol{\delta}\left(x_{3}\right)$

## Convolution algebra

Convolution of $f(\mathbf{x})$ and $g(\mathbf{x})$

$$
(f \star g)(\mathbf{x})=\int_{-\infty}^{\infty} f(\mathbf{y}) g(\mathbf{x}-\mathbf{y}) d \mathbf{y}
$$

## Commutative properties

$$
\begin{gathered}
f \star g=g \star f \\
\frac{\partial}{\partial x_{i}}(f \star g)(\mathbf{x})=f \star \frac{\partial g}{\partial x_{i}}=\frac{\partial f}{\partial x_{i}} \star g
\end{gathered}
$$

## Green's function

3D Green's function for wave equation

$$
\begin{aligned}
\left(\frac{\partial^{2}}{\partial t^{2}}-c_{0}^{2} \nabla^{2}\right) g(\mathbf{x}, t \mid \mathbf{y}, \tau) & =\delta(t-\tau) \boldsymbol{\delta}(\mathbf{x}-\mathbf{y}) \\
g(\mathbf{x}, t \mid \mathbf{y}, \tau) & =\frac{\delta\left\{|\mathbf{x}-\mathbf{y}|-c_{0}(t-\tau)\right\}}{4 \pi c_{0}|\mathbf{x}-\mathbf{y}|}
\end{aligned}
$$

## Lighthill's Acoustic Analogy

## Lighthill's equation

$$
\left(\frac{\partial^{2}}{\partial t^{2}}-c_{0}^{2} \nabla^{2}\right) \rho^{\prime}=\frac{\partial^{2} T_{i j}}{\partial x_{i} \partial x_{j}} .
$$

For cold, isentropic, low Mach-number jets, $T_{i j}$ can be approximated as:

$$
T_{i j}=\rho_{0} u_{i} u_{j}
$$



Figure 1: Geometry for edge scattering

Lighthill eight power law Acoustic power,

$$
P_{a} \sim \frac{\rho_{o} d_{j}^{2}}{c_{0}^{5}} u_{j}^{8}
$$

where $d_{j}$ and $u_{j}$ are the jet exit diameter and velocity, respectively.

## Refraction

Snells's law for determining a ray path is

$$
\begin{equation*}
\frac{\sin \theta}{c_{0}}=\text { constant } . \tag{1}
\end{equation*}
$$

## Diffraction

Field scattered by a sharp edge If the incident plane waves is

$$
\begin{equation*}
p_{\mathrm{i}}(\mathbf{x}, t)=P_{\mathrm{inc}} \exp \left(\mathrm{i} \omega t+\mathrm{i} k_{0} x \cos \theta_{0}+\mathrm{i} k_{0} y \sin \theta_{0}\right), \tag{2}
\end{equation*}
$$

then the diffracted pressure is

$$
\begin{equation*}
p_{\mathrm{d}}=P_{\mathrm{inc}}\left(\frac{2}{\pi k_{0} r}\right)^{\frac{1}{2}} \frac{\sin \left(\theta_{0} / 2\right) \sin (\theta / 2)}{\cos \theta+\cos \theta_{0}} \exp \left(-\frac{\mathrm{i} \pi}{4}-\mathrm{i} k_{0} r\right) . \tag{3}
\end{equation*}
$$

## In a cylindrical duct of radius $a$

The pressure field is given by

$$
p^{\prime}(\mathbf{x}, t)=e^{i(\omega t+n \theta)} J_{n}\left(z_{m n} r / a\right)\left(A e^{-i k x_{3}}+B e^{i k x_{3}}\right),
$$

where $z_{m n}$ is the $m$ th zero of $\dot{J}_{n}(z)$ and $k=\left(k_{0}^{2}-z_{m n}^{2} / a^{2}\right)^{1 / 2}$.
For large azimuthal wavenumber, $n$

$$
z_{1 n} \approx n+1.85 n^{1 / 3}
$$

